

Introduction To Quantum Mechanics

The quantum mechanics deals with the study of the microscopic property of the matter. OR it is the branch of Physics which deals with the small particle whose size must be less than 10^{-10} m .

		Speed	For less than $3 \times 10^8\text{ m/s}$	comparable to
			classical Mech.	Relativistic Mech.
			Quantum Mech.	Quantum field Theory
Large than 10^{-10} m				
Less than 10^{-10} m				

classical mechanics can't explain stem of physical phenomena like quantisation of certain physical properties.

- ① quantisation of certain physical properties
- ② Quantum & entanglement
- ③ Principle of uncertainty
- ④ Wave-particle duality

These phenomena are completely explained by quantum mechanics.

Classical mechanics describe the motion of particles, rigid body, fluids etc. And some formulation of the law of electromagnetism, thermodynamics,

geometrical & physical optics which can explain properly by the new corresponding theory.

Many new phenomena observed during the last decade of 19 century & early years of 20 century couldn't be explained within the framework of classical physics. In an attempt some new phenomena like photoelectric effect of classical physics were modified in favor of new revolutionary concept. This process gave birth to the quantum physics.

Due to the birth of quantum physics some legends of scientist gave the various quantum phenomena which are more useful & more predictable in the revolutionary physics. These scientist are

Planck & Meissner
Einstein
de Broglie

Heisenberg
Schrödinger
Bohr
Dirac

The quantum idea was first introduced by Max Planck in the year of 1900 in an attempt to explain the energy distribution in the spectrum of blackbody radiation. This idea was later successfully explained by Einstein & gave the different phenomena of it.

Neil Bohr use the similar quantum concept to formulate the model of hydrogen atom & observe hydrogen spectra successfully. These development leads the following conclusions that are

- ① radiant energy such as electromagnetic radiation which show the dual nature of radiation.
- ② some dynamical properties such as energy, momentum & angular momentum of the discrete value. So but not the ordinary value. hence the electromagnetic radiation further explained by Louis de Broglie which give the wave particle duality which can explain the dual nature of radiation.

Some of the quantum theory which establish therefore in quantum mechanics states that every system is characterized by a wave function which describe by a wave function.

the system completely. The wave func satisfies a partial differential eqⁿ called the Schrödinger eqⁿ.

The wave func satisfies a partial differential eqⁿ which can again mainly be Maxwell who can give the mechanical formulation of quantum theory & this is the case when the quantum field theory advanced like the quantum mechanics. Similarly various scientist get their contribution towards the quantum mechanics which can help to our day to day life.

Photoelectric effect :-

(i) When a light is incident on the metal surface then some -ve electrons will be released in a particular direction.

(ii) These -ve electrons are called as photoelectrons & the process is called as Photoelectric effect.

Laws of Photoelectric effect :-

(i) Photoelectric effect is an instantaneous process.

(ii) Photoelectric current is directly proportional to the intensity of incident light & is independent of frequency.

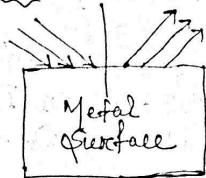
(iii) The Stopping potential & the max velocity of electrons depend upon the frequency of incident light but independent of intensity of incident light.

(iv) The emission of electrons stop below a certain minⁿ frequency known as threshold frequency.

Einstein Photoelectric effect :-

(i) Einstein explained

Photoelectric effect on the basis of Planck's quantum theory.



(ii) According to this theory the radiations composed of small bundles of energy.

(iii) Einstein suggests that these bundles of energy as photons.

(iv) So, according to this theory,

E = hv

Where, h = Planck's constant

v = frequency of radiation.

(v) Again he observed that the electrons are emitted from the metal surface, when they receive some extra energy to overcome the workfunc (W_0) of the metal surface.

(vi) So, according to Einstein,

E = W_0 + K.E

$$\Rightarrow hv = h\nu_0 + \frac{1}{2}mv_{max}^2$$

$$\Rightarrow \frac{1}{2}mv_{max}^2 = hv - h\nu_0 = h(v - \nu_0)$$

where, $V_{\max} = \text{max}^m$ velocity of emitted \bar{e} .
 $m = \text{mass of the electron}$.
 $E = h\nu = \text{energy of the incident Photon}$.

$\omega = h\nu_0 = \text{Worth func.}$

$\nu_0 = \text{threshold frequency.}$

This eqn is known as Einstein's eqn of Photoelectric effect.

Worth func. :-

It is the min^m amount of energy required to pull out an \bar{e} from the surface of metal i.e. $\omega_0 = h\nu_0$.

Compton effect :-

- The effect of change in wavelength on scattering of X-ray photons by electrons is called Compton effect.
- The change in wavelength ($\Delta\lambda$) is called Compton shift.
- The spectrum of the scattered radiation at an angle ' θ ' with the incident direction consist of two components.

- (i) One of the component having wavelength same as that of the incident radiation is called Thomson component.
- (ii) The other component having wavelength greater than the incident wavelength is called Compton component.

Compton also observed that

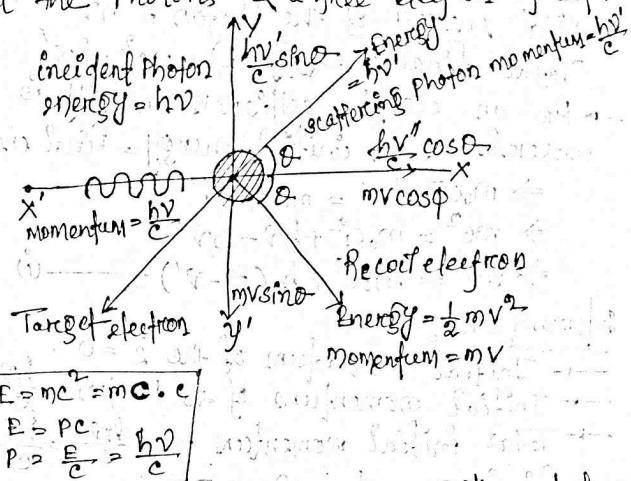
- (i) The difference $\Delta\lambda$ is always, i.e.
- (ii) $\Delta\lambda$ is an increasing func. of the angle ' θ ' bet' the direction of incident & the

Direction of scattering.

(iii) $\Delta\lambda$ depends only on ' θ ' & is independent of the wavelength of the incident radiation & the nature of the scattering material.

Theory :-

According to Compton, the phenomenon of scattering is due to the elastic collision bet' the photons & a free electron at rest.



$$\boxed{E = mc^2 = mc \cdot c \\ E = pc \\ \Rightarrow p = \frac{E}{c} = \frac{h\nu}{c}}$$

(i) In Fig. a X-ray Photon collide with an electron at rest. After collision both the Scattered Photon & the electrons move in different direction.

(ii) Let θ = angle bet' the incident Photon dir. & the Scattered Photon direction.

(iv) ϕ = angle bet' the incident Photon dir. & the recoil electron.

→ ν = frequency of the incident photon.

→ ν' = frequency of the Scattered Photon.

→ m_0 = mass of the rest \bar{e} .

- m = mass of the moving electron
- v = velocity of the electron
- Energy:
- Initial energy of the $e^- = m_0 c^2$
- Initial energy of the photon = $h\nu$
- Total initial energy (E) = $m_0 c^2 + h\nu$
- Final energy of the $e^- = m c^2$
- Final energy of the photon = $h\nu'$
- Total final energy (E') = $m c^2 + h\nu'$
- For an elastic collision energy is conserved. So initial energy = final energy.
- $m_0 c^2 + h\nu = m c^2 + h\nu'$
- $m c^2 = m_0 c^2 + h\nu - h\nu'$
- $m c^2 = m_0 c^2 + h(\nu - \nu')$ — (i)

Momentum:

- Initial momentum of the $e^- = 0$
- Initial momentum of the photon = $\frac{h\nu}{c}$
- Total initial momentum = $0 + \frac{h\nu}{c} = \frac{h\nu}{c}$
- Final momentum of the $e^- = mv \cos \phi$
- Final momentum of the scattered photon = $\frac{h\nu'}{c} \cos \theta$
- Total final momentum = $\frac{h\nu'}{c} \cos \theta + mv \cos \phi$ — (ii)

According to conservation law of momentum =

- Initial momentum = final momentum.
- $m v \cos \phi + \frac{h\nu'}{c} \cos \theta = \frac{h\nu}{c}$
- $m v \cos \phi + \frac{h\nu'}{c} \cos \theta = \frac{h\nu}{c}$
- $m v \cos \phi = \frac{h\nu}{c} - \frac{h\nu'}{c} \cos \theta$
- $m v \cos \phi = \frac{h(\nu - \nu') \cos \theta}{c}$

$$\begin{aligned}
 & \Rightarrow m v \cos \phi = \frac{h(\nu - \nu' \cos \theta)}{c} \\
 & \Rightarrow h(\nu - \nu' \cos \theta) = m v c \cos \phi \quad \text{--- (iii)} \\
 & \text{The Perpendicular component of momentum can be written as} \\
 & \frac{h\nu'}{c} \sin \theta + (-m v \sin \phi) = 0 \\
 & \Rightarrow m v \sin \phi = \frac{h\nu'}{c} \sin \theta \\
 & \Rightarrow m v \sin \phi = h\nu' \sin \theta \quad \text{--- (iv)} \\
 & \text{Squaring & adding eq (iii) & (iv) we get,} \\
 & m^2 v^2 c^2 \cos^2 \phi + m^2 v^2 c^2 \sin^2 \phi = h^2 (\nu - \nu' \cos \theta)^2 \\
 & + [(\frac{h\nu'}{c} \sin \theta)^2 + h^2 \nu'^2 \sin^2 \theta] \\
 & \Rightarrow m^2 v^2 c^2 = h^2 [\nu^2 + \nu'^2 \cos^2 \theta - 2\nu \nu' \cos \theta + \nu'^2 \sin^2 \theta] \\
 & = h^2 [\nu^2 + \nu'^2 \cos^2 \theta - 2\nu \nu' \cos \theta + \nu'^2 \sin^2 \theta] \\
 & \Rightarrow m^2 v^2 c^2 = h^2 (\nu^2 + \nu'^2 - 2\nu \nu' \cos \theta) \quad \text{--- (v)} \\
 & \text{Squaring eq (i) both sides we get,} \\
 & (mc^2)^2 = [m_0 c^2 + h(\nu - \nu')]^2 \\
 & \Rightarrow m^2 c^4 = m_0^2 c^4 + h^2 (\nu^2 + \nu'^2 - 2\nu \nu' \cos \theta) + 2m_0 c^2 (h\nu - h\nu') \quad \text{--- (vi)} \\
 & \text{Subtracting eq (v) from eq (vi) we get,} \\
 & m^2 c^4 - m^2 v^2 c^2 = m_0^2 c^4 + h^2 (\nu^2 + \nu'^2 - 2\nu \nu') + \\
 & 2m_0 c^2 h(\nu - \nu') - h^2 \nu^2 - h^2 \nu'^2 + \\
 & h^2 \cdot 2\nu \nu' \cos \theta \\
 & \Rightarrow m^2 c^4 - m^2 v^2 c^2 = m_0^2 c^4 + h^2 \nu^2 + h^2 \nu'^2 - 2\nu \nu' h^2 + \\
 & 2m_0 c^2 h(\nu - \nu') - h^2 \nu^2 - h^2 \nu'^2 + h^2 \cdot 2\nu \nu' \cos \theta \\
 & \Rightarrow m^2 c^4 (c^2 - v^2) = m_0^2 c^4 + h^2 (\nu^2 + \nu'^2 - 2\nu \nu' - \nu^2 - \nu'^2 + 2\nu \nu' \cos \theta) \\
 & \Rightarrow m^2 c^4 (c^2 - v^2) = m_0^2 c^4 + h^2 (2\nu \nu' - 2\nu \nu' - \nu^2 - \nu'^2 + 2\nu \nu' \cos \theta) \\
 & \Rightarrow m^2 c^4 (c^2 - v^2) = m_0^2 c^4 + h^2 (2\nu \nu' - 2\nu \nu' - \nu^2 - \nu'^2 + 2\nu \nu' \cos \theta) \\
 & \Rightarrow m^2 c^4 (c^2 - v^2) = m_0^2 c^4 + h^2 (2\nu \nu' - 2\nu \nu' - \nu^2 - \nu'^2 + 2\nu \nu' \cos \theta)
 \end{aligned}$$

$$= m_0 c^4 + h^2 [2vv'(\cos\theta - 2vv')] + 2m_0 c^2 h(v-v')$$

$$= m_0 c^4 + h^2 [2vv'(\cos\theta - 1)] + 2m_0 c^2 h(v-v')$$

$$\Rightarrow m_0^2 c^2 (c^2 - v^2) = m_0^2 c^4 + h^2 [2vv'(\cos\theta - 1)] + 2m_0 c^2 h(v-v')$$

Putting the value of $m = \frac{m_0}{\sqrt{1-v^2/c^2}}$ in the above eqn we get

$$\Rightarrow \frac{m_0^2}{1-v^2/c^2} \cdot c^2 (c^2 - v^2) = m_0^2 c^4 + h^2 [2vv'(\cos\theta - 1)] + 2m_0 c^2 h(v-v')$$

$$\Rightarrow \frac{m_0^2 c^2}{(c^2 - v^2)} \cdot c^2 (c^2 - v^2) = m_0^2 c^4 + h^2 [2vv'(\cos\theta - 1)] + 2m_0 c^2 h(v-v')$$

$$\Rightarrow m_0^2 c^4 = m_0^2 c^4 + h^2 [2vv'(\cos\theta - 1)] + 2m_0 c^2 h(v-v')$$

$$\Rightarrow h^2 [2vv'(\cos\theta - 1)] + 2m_0 c^2 h(v-v') = 0$$

$$\Rightarrow h^2 [2vv'(\cos\theta - 1)] = -2m_0 c^2 h(v-v')$$

$$\Rightarrow h[vv'(\cos\theta - 1)] = -m_0 c^2 (v-v') \quad (\text{fact})$$

$$\Rightarrow hvv'(1-\cos\theta) = -m_0 c^2 (v-v')$$

$$\Rightarrow h(1-\cos\theta) = \frac{m_0 c^2 (v-v')}{vv'}$$

$$\Rightarrow \frac{h(1-\cos\theta)}{m_0 c^2} = \frac{(v-v')}{vv'} \quad (\text{fact})$$

$$\Rightarrow \frac{h}{m_0 c^2} (1-\cos\theta) = \frac{v-v'}{vv'} \quad (\text{fact})$$

$$\Rightarrow \boxed{\frac{h}{m_0 c^2} (1-\cos\theta) = \frac{1}{v'} - \frac{1}{v}} \quad (\text{fact})$$

$$\Rightarrow \boxed{\frac{h}{m_0 c} (1-\cos\theta) = \frac{c}{v'} - \frac{c}{v}} \quad (\text{fact})$$

$$\text{Using the formula } c = \nu\lambda \text{ we get } (\because c = \nu\lambda)$$

$$\Rightarrow \frac{h}{m_0 c} (1-\cos\theta) = \lambda' - \lambda = \Delta\lambda$$

$$\Rightarrow \boxed{\Delta\lambda = \frac{h}{m_0 c} (1-\cos\theta)} \quad (\text{vi})$$

This is the expression for compton shift & from the above eqn it is clear that there is increase in wavelength $\Delta\lambda$ is independent of the wavelength of incident radiations.

So, again $1-\cos\theta = 2\sin^2\frac{\theta}{2}$

$$\Rightarrow \boxed{\Delta\lambda = \frac{h}{m_0 c} \cdot 2\sin^2\frac{\theta}{2}}$$

Case-I :- (when $\theta = 0^\circ$)

$$\Delta\lambda = \frac{h}{m_0 c} \cdot 2 \times 0 = 0$$

No, scattering effect

Case-II :- (when $\theta = 90^\circ$)

$$\Delta\lambda = \frac{h}{m_0 c} \times 2 \times \frac{1}{2} = \frac{h}{m_0 c}$$

$$\Rightarrow \boxed{\lambda_c = \frac{h}{m_0 c}} \quad (\text{it is known as compton length})$$

Case-III :- (when $\theta = 180^\circ$)

$$\Delta\lambda = \frac{h}{m_0 c} \times 2 = \frac{2h}{m_0 c}$$

$\Delta\lambda = 2 \times \text{compton wavelength}$ ($\because \frac{h}{m_0 c} = \text{compton wavelength}$)

It gives the max wavelength of the scattered photon, and also it can be written as

$\Delta\lambda = 2 \times \text{compton wavelength}$

Also if it is the special case of head on elastic collision.

Double scattering is found at 180°

Wavelength for head on elastic collision is double the original wavelength.

It will be discussed in part 2 of the notes.

Let's discuss about the scattered field.

Dual nature of radiation :-

- light exhibits the reflection, refraction, dispersion, diffraction, transmission.
- Photo electric effect & compton effect.
- Among the all phenomenon reflection, refraction, dispersion, diffraction & transmission. Show the wave nature of the light particle & the remaining phenomenon photo electric effect & compton effect which show the particle nature of the light particle.
- So, the wavelength of the light particle can be written as,

$$\lambda = \frac{h}{P}$$

Where, h = Planck's constant.
 P = momentum.

- The expression for energy
- $E = h\nu$
- Where, ν = frequency of the light particle.
- So, from the above as the light show the wave nature along with particle nature but not simultaneously. So this is called the dual nature of radiation.

De-Broglie Hypothesis :-

- The electromagnetic radiation shows the dual nature.
- It exhibits the particle nature along with the wave nature but not simultaneously & acts two different experiment.
- If ν is the frequency of the light particle then we can write energy of the light particle $E = h\nu$.

→ In electromagnetic radiation the particle & the wave can stay in the form of energy packet & the energy of the each packet by Einstein is $E = h\nu$.

→ Then in electromagnetic radiation the frequency & the wavelength show the wave character & the momentum & energy show the particle characteristics.

$$\text{So, } E = h\nu = \frac{h}{2\pi} \times 2\pi\nu = h\nu$$
$$P = h\nu = \frac{h}{2\pi} \times \frac{2\pi}{\lambda} = \frac{h}{\lambda}$$

Here, h = Planck's constant.

→ From the above it is clear that the energy & the momentum can relate to each other by a Planck's constant ' h '.

→ So, Lewis De-Broglie Postulates that, if light particle exhibits the wave nature then the particle such as electron also show the particle nature.

→ Hence, light wave exhibit a particle nature & the particle is also wave characteristics.

→ So, according to De-Broglie if λ is the wavelength of any particle in the wave then, the De-Broglie wavelength

$$\lambda = \frac{h}{P}$$

Where, h = Planck's constant

P = momentum.

→ It also can be derived from the relativistic we have, $E = h\nu$

Also in case of relativistic

$$E = \sqrt{m_0 c^2 + P^2 c^2}$$

As the rest mass of photon

$$m_0 = 0$$

$$\text{So, } E = pc \quad \text{(ii)}$$

Comparing eqn (ii) & (i)

$$h\nu = pc$$

$$\Rightarrow p = \frac{h\nu}{c}$$

Again we have, $c = \lambda f$

$$\text{So, } p = \frac{h\nu}{\lambda f} = \frac{h}{\lambda}$$

$$\Rightarrow \boxed{\lambda = \frac{h}{p}}$$

Different forms of De-Broglie Wave length :-

(i) The De-Broglie Wavelength (General case)

$$\lambda = \frac{h}{p} = \frac{h}{mv}$$

Where, m = mass of electron.

$$v = \text{velocity}$$

(ii) In case of non-relativistic case;

$$E_K = \frac{p^2}{2m}$$

$$\therefore p = \sqrt{2mE_K}$$

$$\text{So, } \lambda = \frac{h}{\sqrt{2mE_K}} \quad (\because \lambda = \frac{h}{p})$$

$$\text{Where } E_K = \text{Kinetic energy}$$

(iii) When a charge particle having a charge ' q ' & potential difference ' V ' then the energy of the charge particle,

$$E_K = qV$$

So, the De-Broglie wavelength

$$\lambda = \frac{h}{\sqrt{2mqV}} = \frac{h}{\sqrt{2m\phi}}$$

(In electrostatic field) $\phi = V$

(iv) Consider a matter in thermal radiation,
So, the De-Broglie wavelength

$$\lambda = \frac{h}{\sqrt{2mE_K}} = \frac{h}{\sqrt{2m \cdot \frac{3}{2}kT}}$$

$$\Rightarrow \boxed{\lambda = \frac{h}{\sqrt{3mkT}}}$$

(v) In case of relativistic effect,

$$m = \frac{m_0}{\sqrt{1-v^2/c^2}}$$

So, the De-Broglie wavelength,

$$\lambda = \frac{h}{p} = \frac{h}{mv}$$

$$\Rightarrow \lambda = \frac{h}{mv/\sqrt{1-v^2/c^2}}$$

$$\Rightarrow \boxed{\lambda = \frac{h(1-v^2/c^2)^{1/2}}{mv}}$$

Schrodinger wave dependent Eq :-

Wave function for a particle having energy ' E ' & momentum ' p ' travelling x -direction.

Given by $\psi(x, t) = A e^{i(Et - px)}$

Differentiating twice w.r.t x $\frac{\partial^2 \psi}{\partial x^2} = \frac{i^2}{\hbar^2} (Et - px)$

$$\frac{\partial \psi}{\partial x} = \frac{\partial}{\partial x} A e^{i(Et - px)} = \left(\frac{ip}{\hbar}\right) A e^{i(Et - px)}$$

Again $\frac{\partial^2 \psi}{\partial x^2} = \frac{-p^2}{\hbar^2} (Et - px)$

$$\frac{\partial^2 \psi}{\partial x^2} = \left(\frac{ip}{\hbar}\right)^2 = \frac{-p^2}{\hbar^2} (Et - px)$$

$$\Rightarrow \frac{\partial^2 \psi}{\partial x^2} = \left(-\frac{p^2}{\hbar^2}\right) \psi$$

$$\Rightarrow \frac{\partial^2 \psi}{\partial x^2} = \frac{-p^2}{\hbar^2} \psi$$

$$\rightarrow P^2\psi = -\frac{\hbar^2}{m} \frac{\partial^2 \psi}{\partial x^2} \quad (i)$$

Differentiating eq (i) w.r.t time

$$\frac{\partial \psi}{\partial t} = -\frac{iE}{\hbar} A e^{-iEt/\hbar} (Et - Px)$$

$$= -\frac{iE}{\hbar} \psi$$

$$= \frac{E}{i\hbar} \psi$$

$$\rightarrow i\hbar \frac{\partial \psi}{\partial t} = E\psi \quad (ii)$$

Total energy of a particle is the sum of K.E & P.E i.e $E = \frac{P^2}{2m} + V$

Hence, V = Potential to which the particle is subjected.

Multiplying ψ on both sides

$$E\psi = \frac{P^2\psi}{2m} + V\psi$$

$\rightarrow i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \frac{\partial^2 \psi}{\partial x^2} + V\psi$

This is Schrödinger time dependent eq in one dimension for a particle moving in a potential V .

In 3 dimensions eq becomes

$$\begin{aligned} & -\frac{\hbar^2}{2m} \left[\frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2} + \frac{\partial^2}{\partial z^2} \right] \psi(x, t) + V\psi(x, t) \\ & = i\hbar \frac{\partial \psi}{\partial t} \\ & -\frac{\hbar^2}{2m} \nabla^2 \psi(x, t) + V\psi(x, t) = i\hbar \frac{\partial \psi}{\partial t} \end{aligned}$$

$$\psi = \frac{A e^{i\frac{px}{\hbar} - i\frac{Et}{\hbar}}}{\sqrt{V}}$$

Schrödinger eq for a free particle

for a free particle P.E V is independent of position. Since there is no external force. Hence for a free particle $V=0$.

Schrödinger eq in 1 dimension

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi(x, t)}{\partial x^2} = i\hbar \frac{\partial \psi(x, t)}{\partial t}$$

In 3-dimensions

$$-\frac{\hbar^2}{2m} \nabla^2 \psi(x, t) = i\hbar \frac{\partial \psi(x, t)}{\partial t}$$

Schrödinger Time Independent Eq:

When P.E depends on position only & is independent of time we can separate space & time variable & get the time independent Schrödinger eq. So wave func for a particle of energy E & momentum P travelling along x -direction is

$$\psi(x, t) = A e^{i\frac{px}{\hbar} - i\frac{Et}{\hbar}}$$

$$\psi(x, t) = A e^{i\frac{px}{\hbar} + T(t)}$$

$$\text{Again } \psi(x) = A e^{i\frac{px}{\hbar}} \quad T(t) = e^{i\frac{Et}{\hbar}}$$

$$\text{Now } \psi(x, t) = (\psi(x) \otimes T(t)) \quad (ii)$$

Hence $\psi(x)$ varies with time only & $T(t)$ varies with time only.

Generally Time depends Schrödinger eq is

$$-\frac{\hbar^2}{2m} \frac{\partial^2 \psi(x, t)}{\partial x^2} + V\psi(x, t) = i\hbar \frac{\partial \psi(x, t)}{\partial t} \quad (iii)$$

It is time independent eq of Schrödinger eq

Substituting eq(iii) in eq(iv) we get

$$-\frac{\hbar^2}{2m} \frac{\partial^2}{\partial x^2} [\psi(x) T(t)] + V\psi(x) T(t) = i\hbar \frac{\partial}{\partial t} \psi(x) T(t)$$

$$\Rightarrow -\frac{\hbar^2}{2m} T(t) \frac{\partial^2 \psi(x)}{\partial x^2} + V\psi(x) T(t) = i\hbar \psi(x) \frac{\partial T(t)}{\partial t} \quad (\text{iv})$$

Now $\frac{\partial T(t)}{\partial t} = \frac{\partial}{\partial t} [e^{-iEt/\hbar}] = -\frac{iE}{\hbar} [e^{-iEt/\hbar}]$

$$= -\frac{iE}{\hbar} T(t) \quad (\text{v})$$

Substituting eq(v) in eq(iv) we get

$$-\frac{\hbar^2}{2m} T(t) \frac{\partial^2 \psi(x)}{\partial x^2} + V\psi(x) T(t) = \frac{i\hbar}{\hbar} \psi(x) \left[-\frac{iE}{\hbar} T(t) \right]$$

Dividing through out by $T(t)$

$$\Rightarrow -\frac{\hbar^2}{2m} \frac{\partial^2 \psi(x)}{\partial x^2} + V\psi(x) = \frac{E}{\hbar} \psi(x)$$

$$\Rightarrow \frac{\partial^2 \psi(x)}{\partial x^2} + \frac{2m}{\hbar^2} (E-V) \psi(x) = 0$$

This is Schrodinger time independent eqⁿ in 1-dimension.

In 3-dimension $\nabla^2 \psi(x) + \frac{2m}{\hbar^2} (E-V) \psi(x) = 0$

Solⁿ of Schrodinger wave Eqⁿ:

We know wave func can be written as

$$\psi(x,t) = \psi(x) T(t)$$

In this func $\psi(x)$ depends on potential $V(x)$ & can be obtained by solving the Schrodinger time independent eqⁿ.

The func $T(t)$ is same for all potential & can be determined as follows —

$$T(t) = e^{-iEt/\hbar}$$

$$\frac{\partial T(t)}{\partial t} = -\frac{iE}{\hbar} T(t)$$

Since t is a func of time only we can replace the partial derivative by a complete derivative & write the above eqⁿ as —

$$\frac{\partial T}{\partial t} = -\frac{iE}{\hbar} T$$

$$\Rightarrow \frac{dT}{T} = -\frac{iE}{\hbar} dt$$

Integrating both sides.

$$\ln T = -\frac{iE}{\hbar} t$$

$$\therefore T(t) = e^{-iEt/\hbar}$$

As general wave func is

$$\psi(x,t) = \psi(x) T(t)$$

We can write

$$\psi(x,t) = \psi(x) e^{-iEt/\hbar}$$

This is the solⁿ of Schrodinger wave eqⁿ & we represents the quantum wave eqⁿ.

Solⁿ of Schrodinger eqⁿ

$$\psi(x,t) = \psi(x) e^{-iEt/\hbar}$$

According to maxborn probability density i.e. Probability per unit length of finding the particle in the quantum state $\psi(x,t)$ is given by

$$P(x,t) = \psi^*(x,t) \cdot \psi(x,t)$$

$$= \psi^*(x) e^{iEt/\hbar} \cdot \psi(x) e^{-iEt/\hbar}$$

$$\Rightarrow P(x,t) = \psi^*(x) \psi(x)$$

From the above expression we draw the infinite probability density is independent of time i.e. while these are called stationary state. So $\psi(x,t)$ of a Schrödinger eqⁿ represent as stationary state.

Wave func & its Physical Significance

Wave func ψ is a complex mathematical entity whose variation w.r.t space & time will give raise to matter wave & it also satisfy Schrödinger eqⁿ.

Wave func ψ has no direct physical significance how ever a statistical interpretation of physical significance of the wave func was given by Max Born.

According to Max Born →

- (i) Wave func are mathematical representation of particles which contains all the information required for the probabilistic description of the particle & is generally a complex no.
- (ii) Wave func ψ is called the probability amplitude so that the product of ψ with its complex conjugate gives the probability of particle
- (iii) So probability density is a real & positive no. & hence is an observable quantity.
- (iv) Further, the value of $|\psi|^2$ is there is the probability of finding the particle at a given position & time.

$$(\text{Ans}) \psi^* \psi = 1$$

(v) The Probability that a particle may be found in the region between x & $x+dx$ at a time 't' is given by —

$$P(x,t) dx = \psi^*(x,t) \psi(x,t) dx$$

$$= |\psi(x)|^2 dx$$

Properties of Wave func

The Probability Interpretation of wave func requires that $\psi(x)$ must fulfill following general condition —

- (i) Wave func ψ & its order derivative $\frac{d\psi}{dx}$ must be finite everywhere.
- (ii) This will assure that $|\psi|^2$ which is equal to finding the particle at position 'x' is finite.
- (iii) Wave func ψ must be continuous & its 1st order derivative $\frac{d\psi}{dx}$ must have a single value or unique value at every point.
- (iv) Wave func ψ must be continuous everywhere.
- (v) Wave func ψ must be vanished at infinite $\psi(x) \rightarrow 0$ when $x \rightarrow \pm \infty$.
In other word $\int_{-\infty}^{\infty} \psi^* \psi dx$ should be infinite

Normalisation of a wave func

For a dimensional system probability of finding the particle in a particular volume element d^3 .

$$dP = \psi \psi^* d^3$$

$$dP = |\psi(x,t)|^2 d^3$$

Hence the total probability in all this space is given by

$$P = \int |\Psi(x, t)|^2 dx$$

Normalisation means that the integral of the square of a func should be equal to one i.e. Probability of finding the particle over a space must be equal to one. Hence cond of normalisation of a wave func in one dimension can be written as-

$$\int_{-\infty}^{\infty} |\Psi(x, t)|^2 dx = \int_{-\infty}^{\infty} \Psi^* \Psi dx = 1$$

Any wave func which satisfies the above eq' is called normalisation.

In order to normalise a given wave func $\Psi(x, t)$ we multiply with a normalisation constant A because if Ψ is a sol'

Schrodinger wave eq' then $A\Psi$ is also a sol'

Normalisation cond' becomes

$$\int_{-\infty}^{\infty} (A\Psi)^* (A\Psi) dx = A^2 \int_{-\infty}^{\infty} \Psi^* \Psi dx = 1$$

$$\Rightarrow A^2 = \frac{1}{\int_{-\infty}^{\infty} \Psi^* \Psi dx}$$

Normalised Schrodinger eq' is

$$i\hbar \frac{\partial \Psi}{\partial t} + V\Psi = \frac{-\hbar^2}{2m} \nabla^2 \Psi$$

Probability current density

The Probabilistic interpretation of a wave func Ψ requires that the probability must be a conserve quantity i.e. if the Probability of finding the particle in same bounded region of space decreases in time then the Probability of finding it outside the region must increase by same amount.

We will write the eq' of continuous defining conservation law of Probability which will called as Probability current density.

of P.C. is the Probability of finding the particle at a Particular point in space then the Probability finding the particle in a region of space of the volume V bounded the surface 'S' is given by -

$$P = \int_V P(x) dx = \int_S \Psi^* \Psi dt$$

Where $dx = dy dz$ volume element
In order to find the flow of probability we must know the change in the probability with time with in the region.

So differentiating eq'(i) w.r.t 't'

$$\frac{dP}{dt} = \frac{d}{dt} \int_V (\Psi^* \Psi) dx = \int_V (\Psi \frac{\partial \Psi^*}{\partial t} + \Psi^* \frac{\partial \Psi}{\partial t}) dx$$

Time dependent Schrodinger eq'

$$-\frac{\hbar^2}{2m} \nabla^2 \Psi + V\Psi = i\hbar \frac{\partial \Psi}{\partial t}$$

Complex conjugate of its eq' is

$$-\frac{\hbar^2}{2m} \nabla^2 \Psi^* + V\Psi^* = i\hbar \frac{\partial \Psi^*}{\partial t}$$

Multiplying eq(iii) by Ψ^* & eq(iv) by Ψ
then Substracting eq(iii) & eq(iv), we get

$$-\frac{\hbar^2}{2m} \Psi^* \nabla^2 \Psi + V \Psi \Psi^* = 2\hbar \Psi^* \frac{\partial \Psi}{\partial t} \quad \text{(v)}$$

$$-\frac{\hbar^2}{2m} \Psi \nabla^2 \Psi^* + V \Psi \Psi^* = -2\hbar \Psi \frac{\partial \Psi^*}{\partial t} \quad \text{(vi)}$$

$$-\frac{\hbar^2}{2m} \Psi^* \nabla^2 \Psi^* + V \Psi \Psi^* = \frac{\hbar^2}{2m} \Psi^* \nabla^2 \Psi - V \Psi \Psi^* \quad \text{(vii)}$$

$$\Rightarrow -2\hbar \Psi \frac{\partial \Psi^*}{\partial t} = -2\hbar \Psi^* \frac{\partial \Psi}{\partial t}$$

$$\Rightarrow -\frac{\hbar^2}{2m} \Psi^* \nabla^2 \Psi^* + \frac{\hbar^2}{2m} \nabla^2 \Psi = -2\hbar \Psi \frac{\partial \Psi^*}{\partial t} - 2\hbar \Psi^* \frac{\partial \Psi}{\partial t}$$

$$\Rightarrow \frac{\hbar^2}{2m} (\Psi^* \nabla^2 \Psi - \Psi \nabla^2 \Psi^*) = -2\hbar (\Psi \frac{\partial \Psi^*}{\partial t} + \Psi^* \frac{\partial \Psi}{\partial t})$$

$$\Rightarrow (\Psi \frac{\partial \Psi^*}{\partial t} + \Psi^* \frac{\partial \Psi}{\partial t}) = -\frac{\hbar}{2m} (\Psi^* \nabla^2 \Psi - \Psi \nabla^2 \Psi^*) \quad \text{(v)}$$

Substituting eq(v) in eq(vii) we get

$$\frac{dP}{dt} = -\frac{\hbar}{2m} \int (\Psi^* \nabla^2 \Psi - \Psi \nabla^2 \Psi^*) dV$$

This integral can be transformed into surface integral over the surface 'S' & write

$$\frac{dP}{dt} = -\frac{\hbar}{2im} \oint (\Psi^* \nabla \Psi - \Psi \nabla \Psi^*) dS \quad \text{(vi)}$$

Let's therefore J called Probability current density as —

$$J = \frac{\hbar}{2im} (\Psi^* \nabla \Psi - \Psi \nabla \Psi^*)$$

So eq(vi) reduces to

$$\frac{dP}{dt} = - \oint J dS$$

Using Gauss divergence theorem

$$\frac{dP}{dt} = - \nabla \cdot J$$

$$\Rightarrow \boxed{\frac{dP}{dt} + \nabla \cdot J = 0}$$

→ This eq represents Probability density & shows that the rate of change of Probability of a particle inside a Surface is equal to rate of decrease of Probability current to the surface S. Hence this represents conservation of Probability finding or not finding in a Specified volume.

Operator & Eigen Values:

An operator can be defined as an instruction or mathematical rule which can be used to change by a given function into another function if \hat{A} is an operator, Ψ = wave func then $\hat{A}\Psi = a\Psi$

Where a = Eigen value

Proof $\hat{A}\Psi = \text{Eigen func.}$

Observables & Operators:

Observables are Physical qty such as energy & momentum which gain

Expectation Value:

In quantum mechanics every physical parameters of observable has an inherent uncertainty. General measurement of a parameter of a system under certain conditions doesn't if the same value always even as values of all the measured quantity is called expectation value

of a Physical Parameter α is measured. General rule says that if value x_i appears n_i times in appearance & time & x_n appears n_n times then expectation value of observable α is

$$\langle \alpha \rangle = \frac{\sum_{i=1}^n n_i x_i}{\sum_{i=1}^n n_i} = \frac{\sum_{i=1}^n n_i x_i}{N}$$

Hence $\frac{n_i}{N} = p_i$ = Probability of getting the value x_i .

$$\text{Therefore } \langle \alpha \rangle = \sum_{i=1}^n x_i; \frac{n_i}{N} = \sum_{i=1}^n x_i p_i$$

If the variable α is continuous \leq changes to integration -

$$\langle \alpha \rangle = \int x p(x) dx$$

According to Max Born Probability density p associates with a particle described by the wave func $\psi(x,t)$ is given by -

$$p(x,t) = \psi^*(x,t) \psi(x,t)$$

$$\text{Therefore } \langle \alpha \rangle = \int_{-\infty}^{\infty} x \psi^*(x,t) \psi(x,t) dx$$

$$\text{or } \langle \alpha \rangle = \int_{-\infty}^{\infty} x \psi^* \alpha \psi(x,t) dx$$

so in general expectation value of any observable α is given by

$$\langle \alpha \rangle = \int_{-\infty}^{\infty} \psi^*(x,t) \hat{A} \psi(x,t) dx$$

Where \hat{A} quantens mechanical operator corresponding to the observable "a" and

$$\langle \alpha \rangle = \langle \psi | \hat{A} | \psi \rangle \text{ (Dirac notation)}$$

$$\langle \psi | \rightarrow \text{bra veefor}$$

14) \rightarrow Ket veefor
so expectation value & component of momentum is written as -

$$\langle P_x \rangle = \int_{-\infty}^{\infty} \psi^*(x,t) P_x \psi(x,t) dx$$

$$\text{Since } P_x = (-i\hbar) \frac{\partial}{\partial x}$$

$$\langle P_x \rangle = -i\hbar \int_{-\infty}^{\infty} \psi^*(x,t) \frac{\partial}{\partial x} \psi(x,t) dx$$

$$\langle P_x \rangle = -i\hbar \int_{-\infty}^{\infty} \psi^*(x,t) \Delta \psi(x,t) dx$$

Similarly expectation value of energy is

$$\langle E \rangle = i\hbar \int_{-\infty}^{\infty} \psi^*(x,t) \frac{\partial}{\partial t} \psi(x,t) dx$$